

Derivation of the dipole black ring solutions

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Completely integrable sector in 5D Einstein-Maxwell gravity

$$\begin{aligned} R_{\mu\nu} &= \frac{1}{2} \left(F_{\mu\lambda} F_{\nu}^{\lambda} - \frac{1}{6} F_{\sigma\lambda} F^{\sigma\lambda} g_{\mu\nu} \right), \\ \nabla_{\mu} F^{\mu\nu} &= 0. \end{aligned} \quad (1)$$

Spacetime symmetries: Three commuting Killing vectors $T = \partial/\partial t$, $K_1 = \partial/\partial X$, $K_2 = \partial/\partial Y$.

K_2 -hypersurface orthogonal

Metric:

$$ds^2 = e^{2u} dY^2 + e^{-u} h_{ij} dx^i dx^j \quad (2)$$

h_{ij} - 4D Lorentz metric, $h_{ij} = h_{ij}(x^k)$

Maxwell 2-form :

$$F = dA_Y \wedge dY, \quad A_Y = A_Y(x^k) \quad (3)$$

Dimensionally reduced equations

$$\begin{aligned}
\mathcal{D}_i \mathcal{D}^i u &= -\frac{1}{3} e^{-2u} h^{ij} \mathcal{D}_i A_Y \mathcal{D}_j A_Y, \\
\mathcal{D}_i \left(e^{-2u} \mathcal{D}^i A_Y \right) &= 0, \\
R(h)_{ij} &= \frac{3}{2} \partial_i u \partial_j u + \frac{1}{2} e^{-2u} \partial_i A_Y \partial_j A_Y.
\end{aligned} \tag{4}$$

$$M_1 = \begin{pmatrix} e^u + \frac{1}{3} e^{-u} A_Y^2 & \frac{1}{\sqrt{3}} e^{-u} A_Y \\ \frac{1}{\sqrt{3}} e^{-u} A_Y & e^{-u} \end{pmatrix} \tag{5}$$

$$\mathcal{D}_i \left[\mathcal{D}^i M_1 M_1^{-1} \right] = 0, \tag{6}$$

$$R_{ij}(h) = -\frac{3}{4} \text{Tr} \left[\partial_i M_1 \partial_j M_1^{-1} \right]. \tag{7}$$

Symmetries of the reduced equations:

$$h_{ij} \rightarrow h_{ij}, \quad M_1 \rightarrow G M_1 G^T, \quad G \in SL(2, R) \tag{8}$$

Further reduction along T and K_1

Metric in canonical coordinates

$$h_{ij}dx^i dx^j = -e^{2U} (dt + \mathcal{A}dX)^2 + e^{-2U} \rho^2 dX^2 + e^{-2U} e^{2\Gamma} (d\rho^2 + dz^2) \quad (9)$$

Twist potential

$$\omega = -\frac{1}{2} \star (h) (T \wedge dT) \rightarrow \omega = df \quad (10)$$

$$\partial_\rho f = -\frac{1}{2} \frac{e^{4U}}{\rho} \partial_z \mathcal{A}, \quad (11)$$

$$\partial_z f = \frac{1}{2} \frac{e^{4U}}{\rho} \partial_\rho \mathcal{A}. \quad (12)$$

$$M_2 = \begin{pmatrix} e^{2U} + 4f^2 e^{-2U} & 2f e^{-2U} \\ 2f e^{-2U} & e^{-2U} \end{pmatrix} \quad (13)$$

2D reduced equations

$$\begin{aligned}\partial_\rho \left(\rho \partial_\rho M_1 M_1^{-1} \right) + \partial_z \left(\rho \partial_z M_1 M_1^{-1} \right) &= 0, \\ \partial_\rho \left(\rho \partial_\rho M_2 M_2^{-1} \right) + \partial_z \left(\rho \partial_z M_2 M_2^{-1} \right) &= 0, \\ \rho^{-1} \partial_\rho \Gamma &= -\frac{1}{8} \left[\text{Tr} \left(\partial_\rho M_2 \partial_\rho M_2^{-1} \right) - \text{Tr} \left(\partial_z M_2 \partial_z M_2^{-1} \right) \right] \\ &\quad - \frac{3}{8} \left[\text{Tr} \left(\partial_\rho M_1 \partial_\rho M_1^{-1} \right) - \text{Tr} \left(\partial_z M_1 \partial_z M_1^{-1} \right) \right], \\ \rho^{-1} \partial_z \Gamma &= -\frac{1}{4} \text{Tr} \left(\partial_\rho M_2 \partial_z M_2^{-1} \right) \\ &\quad - \frac{3}{4} \text{Tr} \left(\partial_\rho M_1 \partial_z M_1^{-1} \right).\end{aligned}$$

The sector of Einstein-Maxwell gravity under consideration is completely integrable.

Solution generating method

Let us consider two solutions $M_1 = M^{(1)}$ and $M_2 = M^{(2)}$ of the σ -model equations

$$\partial_\rho \left(\rho \partial_\rho M M^{-1} \right) + \partial_z \left(\rho \partial_z M M^{-1} \right) = 0. \quad (14)$$

In addition let us denote by $\gamma^{(i)}$ the solution of the system

$$\rho^{-1} \partial_z \gamma^{(i)} = -\frac{1}{4} \text{Tr} \left(\partial_\rho M^{(i)} \partial_z M^{(i)-1} \right), \quad (15)$$

$$\rho^{-1} \partial_\rho \gamma^{(i)} = -\frac{1}{8} \left[\text{Tr} \left(\partial_\rho M^{(i)} \partial_\rho M^{(i)-1} \right) - \text{Tr} \left(\partial_z M^{(i)} \partial_z M^{(i)-1} \right) \right]. \quad (16)$$

Then we find for the metric function Γ

$$\Gamma = \gamma^{(2)} + 3\gamma^{(1)}. \quad (17)$$

From a practical point of view it is more convenient to associate the σ -model solutions $M^{(i)}$ with solutions of the vacuum Einstein equations*

$$ds_{E^{(i)}}^2 = e^{2u_E^{(i)}} dY^2 + e^{-u_E^{(i)}} \left[-e^{2U_E^{(i)}} \left(dt + \mathcal{A}_E^{(i)} dX \right)^2 + e^{-2U_E^{(i)}} \rho^2 dX^2 + e^{-2U_E^{(i)}} e^{2\Gamma_E^{(i)}} (d\rho^2 + dz^2) \right],$$

which correspond to the matrixes

$$M^{(i)} = \begin{pmatrix} e^{2U_E^{(i)}} + 4 \left(f_E^{(i)} \right)^2 e^{-2U_E^{(i)}} & 2f_E^{(i)} e^{-2U_E^{(i)}} \\ 2f_E^{(i)} e^{-2U_E^{(i)}} & e^{-2U_E^{(i)}} \end{pmatrix}.$$

*From now on all quantities with subscript or superscript "E" correspond to the vacuum case.

The metric function $\Gamma_E^{(i)}$ for the vacuum Einstein equations can be found from the equations of Γ by setting $A_Y = 0$ in the matrix M_1 . So we obtain

$$\Gamma_E^{(i)} = \gamma^{(i)} + \Omega_E^{(i)} \quad (18)$$

where $\Omega_E^{(i)}$ is a solution to the system

$$\rho^{-1} \partial_\rho \Omega_E^{(i)} = \frac{3}{4} \left[\left(\partial_\rho u_E^{(i)} \right)^2 - \left(\partial_z u_E^{(i)} \right)^2 \right] \quad (19)$$

$$\rho^{-1} \partial_z \Omega_E^{(i)} = \frac{3}{2} \partial_\rho u_E^{(i)} \partial_z u_E^{(i)}. \quad (20)$$

We then find from (17) and (18) that

$$\Gamma = \Gamma_E^{(2)} - \Omega_E^{(2)} + 3 \left[\Gamma_E^{(1)} - \Omega_E^{(1)} \right]. \quad (21)$$

Comparing the matrixes M_1 and $M^{(1)}$ we obtain

$$e^{2u} = e^{4U_E^{(1)}}, \quad (22)$$

$$A_Y = 2\sqrt{3}f_E^{(1)}, \quad (23)$$

where $f_E^{(i)}$ satisfies

$$\partial_\rho f_E^{(i)} = -\frac{1}{2} \frac{e^{4U_E^{(i)}}}{\rho} \partial_z \mathcal{A}_E^{(i)}, \quad (24)$$

$$\partial_z f_E^{(i)} = \frac{1}{2} \frac{e^{4U_E^{(i)}}}{\rho} \partial_\rho \mathcal{A}_E^{(i)}. \quad (25)$$

Once having the metric function $e^{2u} = g_{YY}$ we can write the EM metric

$$\begin{aligned}
ds^2 = & e^{4U_E^{(1)}} dY^2 + e^{-2U_E^{(1)}} \left[-e^{2U_E^{(2)}} \left(dt + \mathcal{A}_E^{(2)} dX \right)^2 \right. \\
& \left. + e^{-2U_E^{(2)}} \rho^2 dX^2 \right. \\
& \left. + \left(\frac{e^{2\Gamma_E^{(1)}}}{e^{2\Omega_E^{(1)}} + \frac{2}{3}\Omega_E^{(2)}} \right)^3 e^{-2U_E^{(2)}} e^{2\Gamma_E^{(2)}} (d\rho^2 + dz^2) \right].
\end{aligned}$$

Summarizing, we obtain the following result

Proposition. [hep-th/0602116] *Let us consider two solutions of the vacuum 5D Einstein equations*

$$ds_{E(i)}^2 = g_{YY}^{E(i)} dY^2 + g_{00}^{E(i)} \left(dt + \mathcal{A}_E^{(i)} dX \right)^2 + \tilde{g}_{XX}^{E(i)} dX^2 + g_{\rho\rho}^{E(i)} (d\rho^2 + dz^2)$$

Then the following give a solution to the 5D EM equations

$$ds^2 = \left[|g_{00}^{E(1)}| \sqrt{g_{YY}^{E(1)}} \right]^2 dY^2 + \left[\frac{\sqrt{g_{YY}^{E(2)}}}{|g_{00}^{E(1)}| \sqrt{g_{YY}^{E(1)}}} \right] \times \left[g_{00}^{E(2)} \left(dt + \mathcal{A}_E^{(2)} dX \right)^2 + \tilde{g}_{XX}^{E(2)} dX^2 + \left(\frac{|g_{00}^{E(1)}| g_{YY}^{E(1)} g_{\rho\rho}^{E(1)}}{e^{2\Omega_E^{(1)} + \frac{2}{3}\Omega_E^{(2)}}} \right)^3 g_{\rho\rho}^{E(2)} (d\rho^2 + dz^2) \right],$$

$$A_Y = \pm 2\sqrt{3} f_E^{(1)},$$

where $f_E^{(1)}$ is a solution to the system

$$\begin{aligned}\partial_\rho f_E^{(1)} &= -\frac{1}{2} \frac{(g_{00}^{E(1)})^2 g_{YY}^{E(1)}}{\rho} \partial_z \mathcal{A}_E^{(1)}, \\ \partial_z f_E^{(1)} &= \frac{1}{2} \frac{(g_{00}^{E(1)})^2 g_{YY}^{E(1)}}{\rho} \partial_\rho \mathcal{A}_E^{(1)},\end{aligned}$$

and $\Omega_E^{(i)}$ satisfy

$$\begin{aligned}\rho^{-1} \partial_\rho \Omega_E^{(i)} &= \frac{3}{16} \left[\left(\partial_\rho \ln \left(g_{YY}^{E(i)} \right) \right)^2 - \left(\partial_z \ln \left(g_{YY}^{E(i)} \right) \right)^2 \right], \\ \rho^{-1} \partial_z \Omega_E^{(i)} &= \frac{3}{8} \partial_\rho \ln \left(g_{YY}^{E(i)} \right) \partial_z \ln \left(g_{YY}^{E(i)} \right).\end{aligned}$$

Explicit derivation of the dipole black ring solutions

$$\{\textit{neutral black ring}\} + \{\textit{neutral black ring}\} \rightarrow \{\textit{EM dipole black ring}\}$$

The first solution is with parameters $\{\lambda_1, \nu, \mathcal{R}\}$ while the second is parameterized by $\{\lambda_2, \nu, \mathcal{R}\}$.

The derivation of the dipole black rings with dilaton can be found in hep-th/0604140 and hep-th/0607101.

some unresolved problems

1. $\{\textit{neutral black ring}\} + \{\textit{neutral black hole}\} \rightarrow$

$\{\text{???\}$

2. $\{\textit{neutral black hole}\} + \{\textit{black hole}\} \rightarrow$

$\{\text{???\}$

3. $\{\textit{appropriate solution}\} + \{\textit{black hole/ring}\} \rightarrow$

$\{\textit{black soluton ???}\}$

4. Are there dipole black ring solutions with rotation in ϕ -direction? If YES, could they be generated via the presented solution generating method?

$\{\text{black ring with rotation in } \phi\text{-direction}\} + \{\text{?????}\}$
 $= \{\text{dipole black ring with rotation in } \phi\text{-direction}\}$

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